Developing the Prototype Schwarzschild-Couder Telescope for the Cherenkov Telescope Array

Brent Mode University of Wisconsin, Madison 13 October 2020

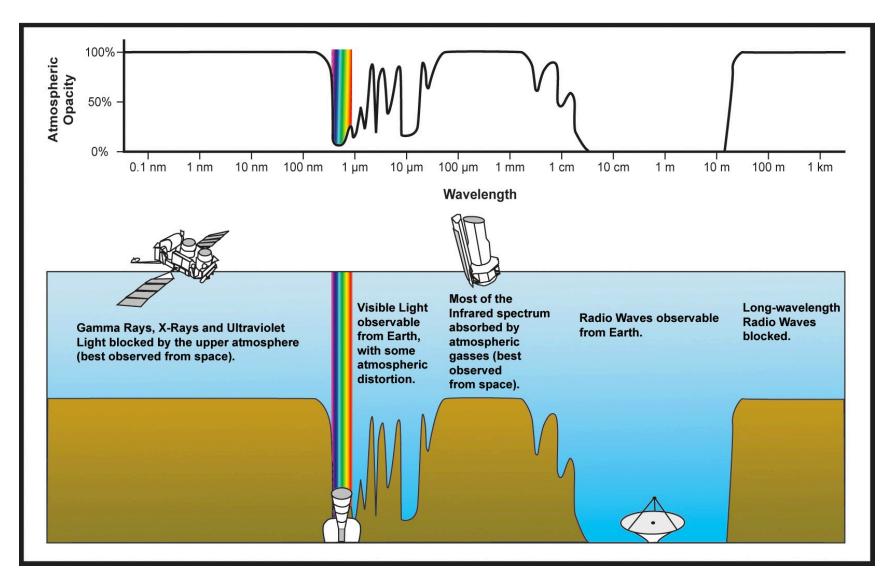
Questions To Be Answered:

- What is a Cherenkov telescope?
- Why would we like to build one of those?
- How do we make the best one yet?

Cta pSCT

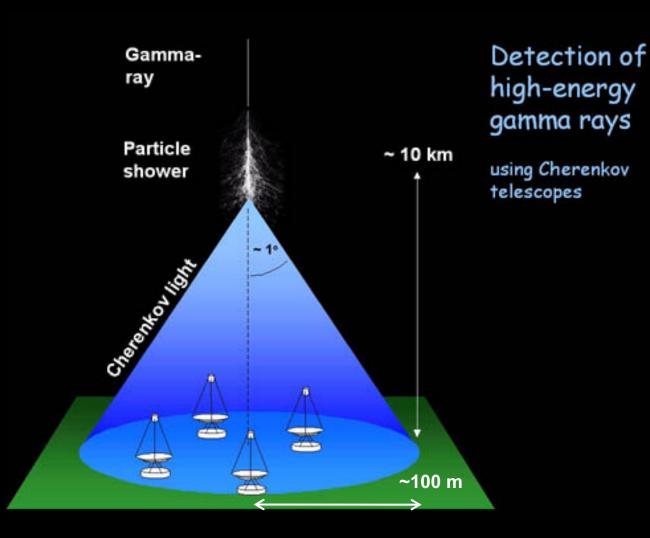
What is a Cherenkov Telescope?

The Atmosphere is Opaque to Gamma Rays



Atmospheric Cherenkov Radiation

- Optical frequency (blue) light
- Very short (few ns) exposure to limit night sky background
- Cherenkov cone very narrow, ~1°:
- $\theta = \arccos \frac{1}{n\beta}$
- 1000-1500 hours per year (dark, good weather)

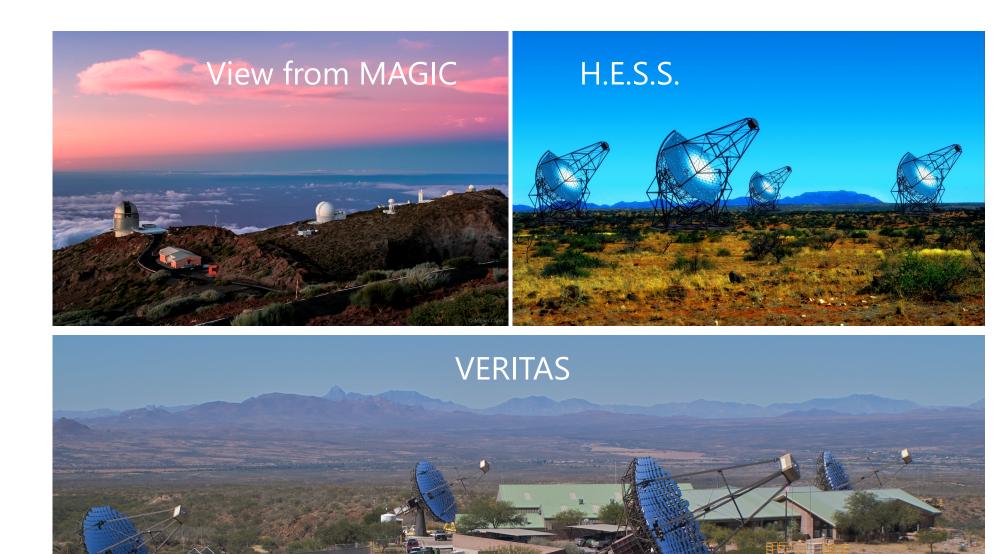


First IACT: Whipple 10 m Telescope at FLWO



- Pioneer imaging atmospheric Cherenkov telescope
- Discovered the first very-high energy (TeV) astronomical sources
 - Crab Nebula: 1985 (Optical), 1989 (UV)
 - Markarian 421 (1992): a nearby blazar
 - Markarian 501 (1997): another nearby blazar

Current Generation of IACTs



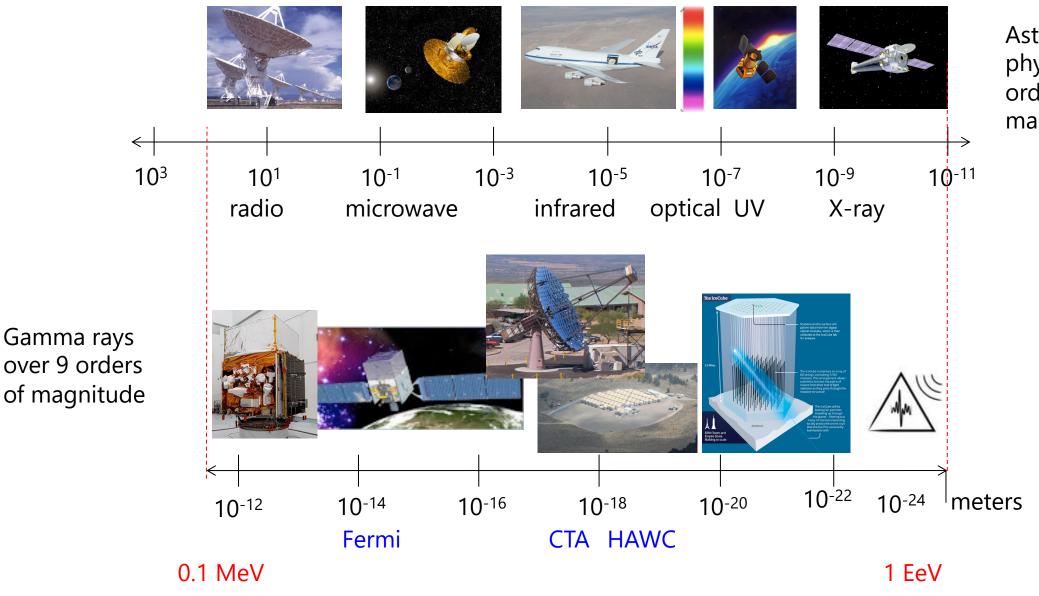
Why Would We Like to Build One of These?

13 October 2020

Brent Mode

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All the Different Colors of Light

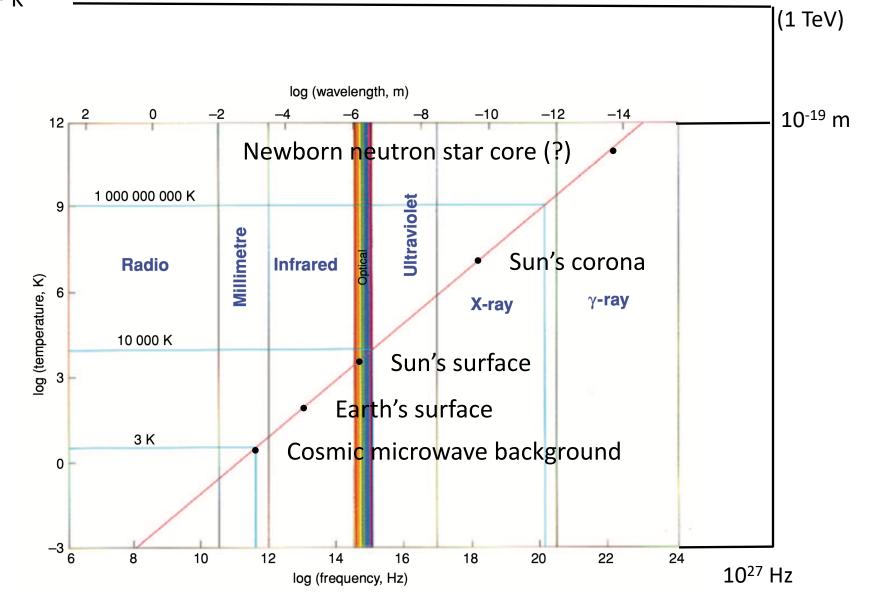


Astroparticle physics over 13 orders of magnitude

over 9 orders of magnitude

The Thermal v. Non-Thermal Universe

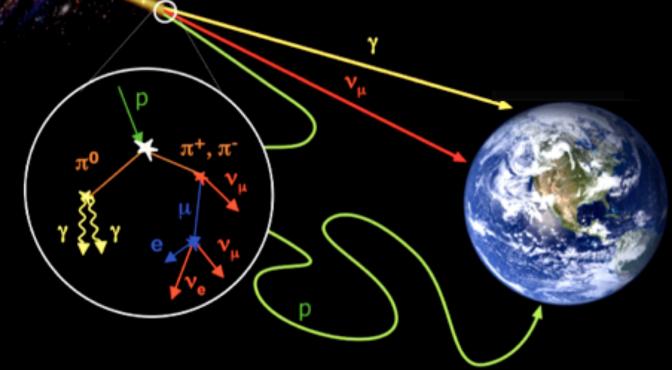
- Black body radiation is responsible for much of the low energy light in the universe
- Even some gamma rays can come from very high energy thermal events
- Most gamma rays will come from non-thermal processes, as the associated black body temperature peaked at 1 TeV is 10 quadrillion K



CTA

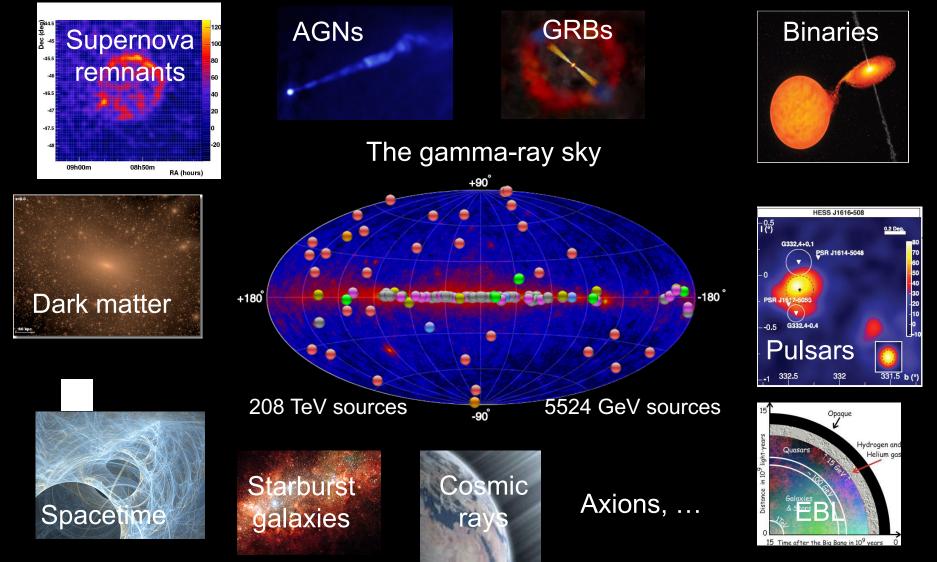
Multi-Messenger Astronomy

Astrophysical beam dump



- Using photons, neutrinos, cosmic rays, and gravitational waves, we can study astrophysical sources and transient objects much more thoroughly than ever before
- Different astrophysical sources emit different particles and at different energies, allowing for multiinstrument, coordinated observations

Physics with TeV Gamma Ray Telescopes



How Do We Make the Best One Yet?

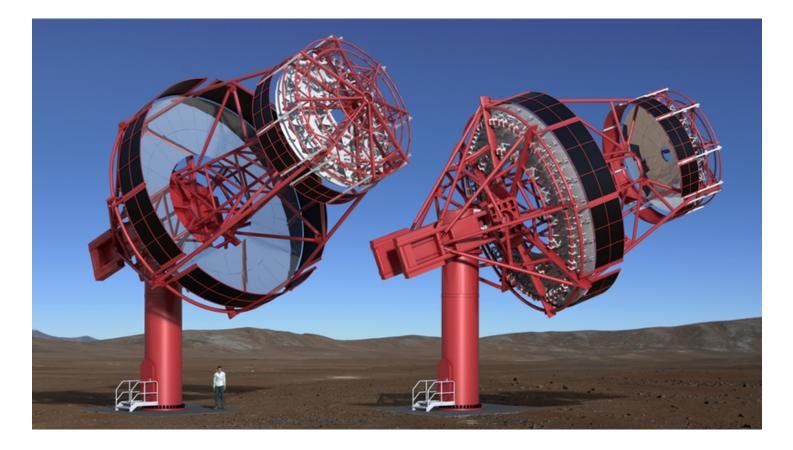
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pSCT: Prototype Schwarzschild-Couder Telescope

Use two mirrors instead of one:

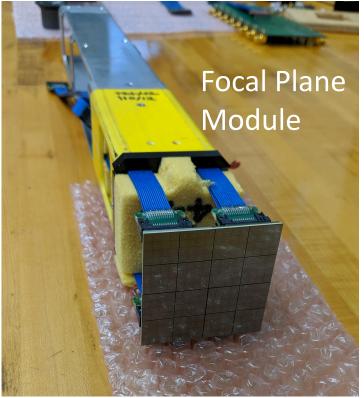
- Advantages:
 - Telescope can be more compact
 - Has wider field of view
 - Better resolution
- Need special technique for aspherical mirror shaping:
 - optimized for maximum resolution and field of view
- Need fast, high-resolution camera:
 - possible through new developments in SiPM and ASIC technology



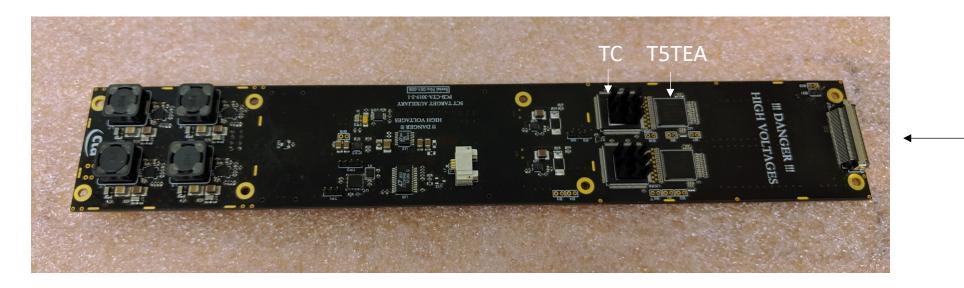
TARGET C and FPM

- TeV Array Readout with GSa/s sampling and Event Trigger – TARGET
- Light detected by SiPMs on the focal plane module
- Telescope can record 1,000,000,000 images per second
 - Needed to image an air shower accurately
- Holds 16 μs of images at once
 - Needed to give enough time to communicate with other telescopes
- UW group has been a strong contributor to the development of camera modules and commissioning of the pSCT camera





TARGET C Module





To Backplane —

Signal Input

Analyzing Data from the pSCT

- New telescope, so we need to understand the kind of data that we are getting
- Need to do image classification to select the useful data
- Then need to clean up that data to do science
- Finally, need to do some science

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Image Classification

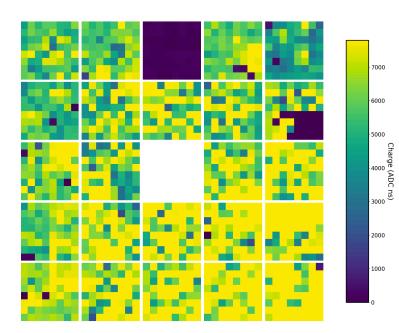
Run 328555

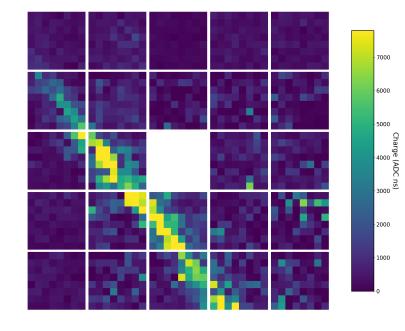
Shower

Event 393

Flasher

Event 2709







Event 112538

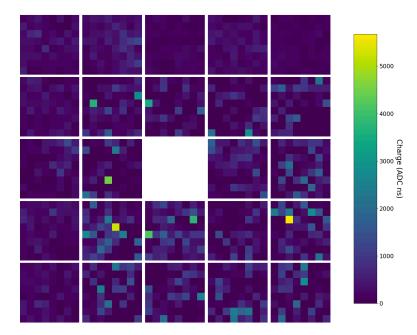
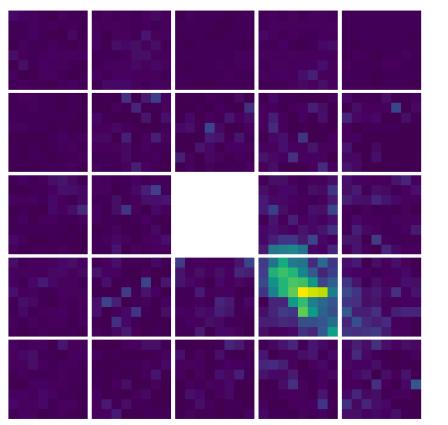


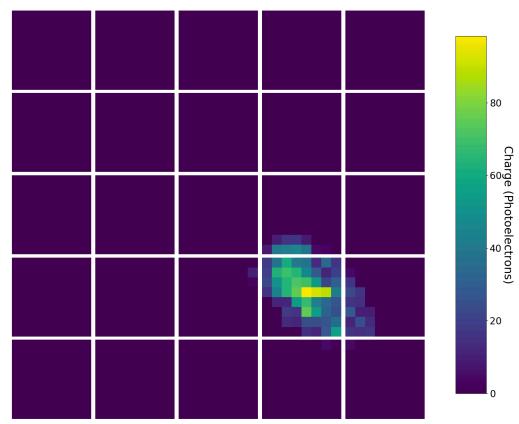
Image Cleaning

Run 328555 Event 1826 2020-01-18 02:56:08 UTC





Run 328555 Event 1826 2020-01-18 02:56:08 UTC



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Detecting the Crab Nebula

- The Crab Nebula is the brightest constant gamma ray source in the sky
- pSCT observed the Crab Nebula from January to February 2020
- Resulted in this nice detection
- Paper in preparation

